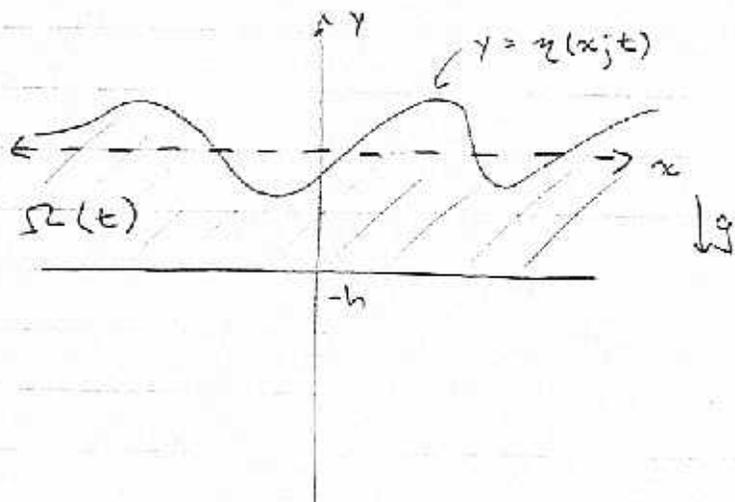


# "How Accurate is KdV Approximation to the Water Wave Equation?"

Fields Seminar  
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The Water Wave Equation As we saw it last week:



If the free surface is a graph over the  $x$ -coordinate, i.e.

$$\Gamma(t) = \text{free surface} = \{(x, \eta(x; t)) \mid x \in \mathbb{R}, t \geq 0\}$$

And if we denote the fluid velocity at position  $(x, y)$  at time  $t$  by

$\underline{u}(x, y; t)$ , then the evolution of the surface is given by:

WE {

$$\left. \begin{aligned} \nabla \cdot \underline{u} &= 0 \\ \nabla \times \underline{u} &= 0 \end{aligned} \right\} \text{ in } \Omega(t) \leftarrow \begin{array}{l} \text{incompressible} \\ \text{irrotational} \end{array}$$

$$u_z(x, -h) = 0 \leftarrow \text{no flow through bottom}$$

$$\partial_z \eta = u_z - u_x \partial_x \eta \quad \text{on free surface} \leftarrow \text{The kinematic boundary condition}$$

$$\partial_t \underline{u} + (\underline{u} \cdot \nabla) \underline{u} = -\nabla p - \begin{pmatrix} 0 \\ g \end{pmatrix} \quad \text{on free surface} \leftarrow \text{Euler's Eqn}$$

## The "Eulerian" Formulation

We can reformulate this as follows. Let us track the position of fluid particles, instead of having a fixed lab frame.

Step 1 Parameterize the free surface as follows

$$\Gamma(t) = \{ (\tilde{X}(\alpha;t), Y(\alpha;t)) \mid \alpha \in \mathbb{R}, t \geq 0 \}$$

Here  $\underbrace{(\tilde{X}(\alpha;t), Y(\alpha;t))}_{\underline{X}(\alpha;t)}$  = Position at time  $t$  of the "particle" which was at  $(\tilde{X}(\alpha;0), Y(\alpha;0))$  at time  $t=0$ .

We will find equations which give the evolution of  $\tilde{X}$  &  $Y$ .

The key change of variables is:

$$\partial_t \underline{X}(\alpha;t) = \underline{u}(\underline{X}(\alpha;t), t)$$

So:

$$\begin{aligned} \partial_t^2 \underline{X} &= \partial_t \underline{u} + (\underline{u} \cdot \nabla) \underline{u} \\ &= -\nabla p - \begin{pmatrix} 0 \\ g \end{pmatrix} \end{aligned}$$

on the surface  
From (W.E.)

or rather

$$\partial_t^2 \underline{X}(\alpha;t) = -\nabla p(\underline{X}(\alpha;t), t) - \begin{pmatrix} 0 \\ g \end{pmatrix}$$



Now the pressure,  $p$ , is constant on the free surface, i.e.

$$p(\underline{x}(\alpha, t), t) = \text{constant}$$

$$\partial_\alpha (p(\underline{x}(\alpha, t), t)) = 0$$

$$\nabla p(\underline{x}(\alpha, t), t) \cdot (\partial_\alpha \tilde{X}(\alpha, t), \partial_\alpha Y(\alpha, t)) = 0$$

So we remove  $\nabla p$  from  $\textcircled{2}$  by taking the inner product of  $\textcircled{2}$  w/  $(\partial_\alpha \tilde{X}, \partial_\alpha Y)$  and we have

$$(\partial_t^2 \tilde{X}, \partial_t^2 Y) \cdot (\partial_\alpha \tilde{X}, \partial_\alpha Y) = -(0, g) \cdot (\partial_\alpha \tilde{X}, \partial_\alpha Y)$$

$$\partial_t^2 \tilde{X} \partial_\alpha \tilde{X} + \partial_t^2 Y \partial_\alpha Y = -g \partial_\alpha Y$$

I am lazy, so  $g = 1$ . For reasons technical, we take

$$\tilde{X}(\alpha, t) = \alpha + X(\alpha, t)$$

we have

$$\partial_t^2 X (1 + \partial_\alpha X) + \partial_\alpha Y (1 + \partial_t^2 Y) = 0$$

Link 1: the kinematic bndry condition reduces in a similar fashion to

$$\partial_t Y = u_2 \leftarrow \text{which is boring}$$

Now notice

$$\begin{aligned} \partial_x u_1 + \partial_y u_2 &= 0 \\ \partial_x u_2 - \partial_y u_1 &= 0 \end{aligned}$$

Are almost C-R eqns (- sign problem)

the <sup>Complex</sup> function  $f(x+iy) = u_1(x,y) - i u_2(x,y)$  is analytic in  $\Omega(t)$ .

Def'n: Given a region  $\Omega$  in  $\mathbb{C}$ , with boundary  $\partial\Omega$  there is a <sup>(linear)</sup> relationship between the real and imaginary components of analytic function on  $\Omega$  when restricted to  $\partial\Omega$ . That is,  $\exists H(\Omega)$  linear such that if  $f(z) = u(x,y) + i v(x,y)$

$$v|_{\partial\Omega} = H(\Omega) u|_{\partial\Omega}$$

Sol

$$u_2 = H(\Omega(t))(-u_1)$$

Called the "Hilbert Transform of  $\Omega$ "

and recalling

$$\begin{aligned} u_2(x, y(x,t)) &= \partial_t Y(x,t) \\ u_1(x, y(x,t)) &= \partial_t X(x,t) \end{aligned}$$

we have

$$\partial_t Y = -H(\Omega(t)) \partial_t X$$

Note:  $\Omega(t)$  is determined by  $(X, Y)$ , so we let

$$K(X(x,t), Y(x,t)) = -H(\Omega(t))$$

Eq 1 If  $\Omega =$  lower half plane,

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then  $H(\Omega) =$  more typical Hilbert transform  $H$

where  $\widehat{Hf}(k) = i \operatorname{sgn}(k) \widehat{f}(k)$  ← lets go on Fourier Transform

Eq 2 If  $\Omega_0 = \{(x, y) \mid x \in \mathbb{R}, -1 \leq y \leq 0\}$  ← unit strip

and if  $f(z)$  is analytic in  $\Omega_0$  and  $\operatorname{Im} f \Big|_{(x, -1)} = 0$

then

$$\widehat{H(\Omega_0)u}(k) = i \tanh(k) \widehat{u}(k)$$

Ch:  $f = u + iv$  analytic  $\Rightarrow v_{xx} + v_{yy} = 0$  in  $\Omega_0$

$$\text{and } \partial_x u = \partial_y v$$

Take F.T. wrt  $x$  coord

$$\leadsto -k^2 \widehat{v} + \partial_y^2 \widehat{v} = 0 \Rightarrow \partial_y^2 \widehat{v}(k, y) = -k^2 \widehat{v}(k, y) \quad \& \quad \widehat{v}(k, -1) = 0$$

$$\widehat{v}(k, y) = A(k) \sinh(k(y+1))$$

$$ik \widehat{u}(k, y) = \partial_y \widehat{v}(k, y) = A(k) k \cosh(k(y+1))$$

$$\leadsto \widehat{u}(k, y) = -i A(k) \cosh(k(y+1))$$

$$\leadsto \frac{\widehat{v}(k, y)}{\widehat{u}(k, y)} = i \tanh(k(y+1))$$

$$\leadsto \boxed{\widehat{v}(k, 0) = i \tanh(k) \widehat{u}(k, 0)}$$



If  $\nu = -1$ , then (under certain conditions)

Thank you  
W. Craig!!

$$K(x, y)u = K_0 u + K_1(x, y)u + S_2(x, y)u$$

where  $K_0 u = \widehat{K_0(k)} \widehat{u}(k)$

$$K_1(x, y)u = [x, K_0] \partial_x u - (y + K_0 \gamma K_0) \partial_y u$$

$$\& \quad S_2(x, y) = \text{Quadratic in } x, y$$

Sol

$$\partial_t^2 X (1 + \partial_x X) + \partial_x Y (1 + \partial_x^2 Y) = 0 \quad (1)$$

$$\partial_t Y = K(X, Y) \partial_x X \quad (2)$$

WW

There are reasons why this formulation is better. For our purpose it's better because things can be proven in this setting that are hard to prove in WE.

Q: Hey, where's  $k \partial V / \partial u$ ?

More Brute force than we saw last week. <sup>But still</sup> (Formal)

Assume long waves, small amplitudes, slowly varying solns

→ take  $0 < \epsilon < 1$

$$\beta = \epsilon \alpha, \quad \tau = \epsilon t \quad ?$$

$$\left. \begin{aligned} X(\alpha, t) &= \epsilon X(\beta, \tau) + \mathcal{O}(\epsilon^3) \\ Y(\alpha, t) &= \epsilon^2 Y_1(\beta, \tau) + \epsilon^4 Y_2(\beta, \tau) + \mathcal{O}(\epsilon^6) \end{aligned} \right\} \text{Does this scaling mystify you?}$$

Notice  $\sim \partial_t \rightsquigarrow \epsilon \partial_\tau \quad \partial_x \rightsquigarrow \epsilon \partial_\beta$

Also Notice we can expand  $\hat{K}_0(k)$  in T.S.

$$-i \tanh(k) = -ik - \frac{1}{3}(ik)^3 - \frac{2}{15}(ik)^5 + \mathcal{O}(k^7)$$

• knowing under F.T. that  $\partial_x \longleftrightarrow ik$

we see

$$K_0 = -\partial_x - \frac{1}{3}\partial_x^3 + \mathcal{O}(\partial_x^5) \quad \left. \begin{array}{l} \text{or in long wave} \\ \text{limit} \end{array} \right\}$$

$$K_0 \sim -\varepsilon \partial_x - \frac{\varepsilon^3}{3}\partial_x^3 + \mathcal{O}(\varepsilon^5)$$

Prop  $\|K_0 f + \partial_x f\|_s \leq C \|\partial_x^3 f\|_s \quad \text{if } f \in H^{s+3}, s \geq 0$

$$\begin{aligned} & \int_{\mathbb{R}^3} (1+k^2)^s |(\hat{K}_0(k) + ik) \hat{f}(k)|^2 dk \\ & \leq \int_{\mathbb{R}^3} (1+k^2)^s C|k|^6 |\hat{f}(k)|^2 dk \\ & \leq C \|\partial_x^3 f\|_s^2 \end{aligned}$$

Cor. If  $f(x) = F(\varepsilon x)$ , then

$$\|K_0 f + \partial_x f\|_s \leq C \varepsilon^{s/2} \|F\|_{s+3}$$

Remark: we get similar estimates for higher approximations of  $K_0$ . We trade smoothness for smallness

Point We can now make long wave approximations to every thing in (WW)

Start:

$$\partial_t \gamma = K(X, Y) \partial_t X$$

$$= (K_0 + \Sigma X, K_0) \partial_x - (\gamma + K_0(\gamma K_0)) \partial_x + S_2(\dots) \partial_t X$$

$$\partial_t \gamma = \epsilon^3 \partial_t \gamma_1 + \epsilon^5 \partial_t \gamma_2 + O(\epsilon^7)$$

$$\text{RHS} = \epsilon^3 (-\partial_{\beta} \partial_t x) + \epsilon^5 (-\frac{1}{3} \partial_{\beta}^3 \partial_t x + \partial_{\beta} x \partial_{\beta} \partial_t x - \gamma_1 \partial_{\beta} \partial_t x)$$

$$\begin{aligned} \text{so } \partial_t \gamma_1 &\approx -\partial_{\beta} \partial_t x \\ \Rightarrow \gamma_1 &= -\partial_{\beta} x \end{aligned} \quad \left. \vphantom{\begin{aligned} \partial_t \gamma_1 \\ \Rightarrow \gamma_1 \end{aligned}} \right\} O(\epsilon^3)$$

+ sim at  $O(\epsilon^5)$

$$\gamma_2 = -\frac{1}{3} \partial_{\beta}^3 x + (\partial_{\beta} x)^2$$

Now we have  $\gamma_1, \gamma_2$  in terms of  $x$ , so sub everything into (WW1)

$$\partial_t^2 \partial_{\beta} x = \partial_{\beta}^3 x + \epsilon^2 \left( \frac{1}{3} \partial_{\beta}^5 x - \frac{3}{2} \partial_{\beta}^2 (\partial_{\beta} x)^2 \right)$$

if  $z = -\partial_{\beta} x$  we get:

$$\partial_t^2 z - \partial_{\beta}^2 z = \epsilon^2 \left( \frac{1}{3} \partial_{\beta}^4 z + \frac{3}{2} \partial_{\beta}^2 (z^2) \right)$$

Boussinesq

To get to KdV, let

$$z = A_r(\underbrace{\beta}_\beta - \varepsilon, \varepsilon^2 t) + A_l(\underbrace{\beta}_\beta + \varepsilon, \varepsilon^2 t)$$

+ we have: (to  $\mathcal{O}(\varepsilon^4)$ )

$$\left. \begin{aligned} 2 \partial_T A_r &= -\frac{1}{3} \partial_{\beta_-}^3 A_r - \frac{3}{2} \partial_{\beta_-} (A_r^2) \\ 2 \partial_T A_l &= \frac{1}{3} \partial_{\beta_+}^3 A_l + \frac{3}{2} \partial_{\beta_+} (A_l^2) \end{aligned} \right\} \text{KdV}$$

Backing up

formally:

$$\begin{aligned} -\partial_x X &= \varepsilon^2 (A_l(\varepsilon(\alpha+t), \varepsilon^2 t) + A_r(\varepsilon(\alpha-t), \varepsilon^2 t)) + \mathcal{O}(\varepsilon^4) \\ Y(\alpha, t) &= \varepsilon^2 (A_l(\varepsilon(\alpha+t), \varepsilon^2 t) + A_r(\varepsilon(\alpha-t), \varepsilon^2 t)) + \mathcal{O}(\varepsilon^4) \end{aligned}$$

where  $A_l, A_r$  satisfy KdV eqns

How do we prove that this is true?

~~Step 1. Rewrite (now) as an evolution equation in a "good" set of variables (preferably things will be quadratic).~~

Point 1 if one examines the linearized eqns, one finds that  $X^{(lin)}$  will grow linearly in time, but  $\partial_x X$  and  $K_0 X \sim -\partial_x X$  do not so we will examine the W.W.E in the following variables

$$Z = K_0 X$$

$$U = \partial_t X$$

and  $Y$

So we have seen formally  $Z \sim \epsilon^2 (A_L + A_R)$   
 $Y \sim \epsilon^2 (A_L + A_R) \sim \epsilon^2 \mathcal{L}^Y$   
 one can check that  $U \sim \epsilon^2 (A_R - A_L) \sim \epsilon^2 \mathcal{L}^U$

So we want to measure the difference between a true sol'n to WW & the approx. Thus we let

$$Z^{(a,t)} = \epsilon^2 \mathcal{L}^Z + \epsilon^{7/2} R^Z(x,t)$$

$$Y^{(a,t)} = \epsilon^2 \mathcal{L}^Y + \epsilon^{7/2} R^Y$$

$$U^{(a,t)} = \epsilon^2 \mathcal{L}^U + \epsilon^{7/2} R^U$$

This Analysis compliments of G. Schneider & C. Wayne

and we want to derive the evolution equations for  $(R^Z, R^Y, R^U)$ . If we can do so in a way which proves  $R^Z, R^Y, R^U$  remain  $\Theta(1)$  for some amount of time, we're okay!

THIS is Wicked Hard to do!

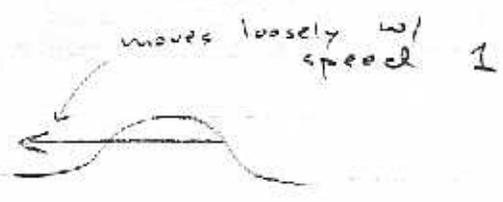
End result you have  $\epsilon_0$  st : $\epsilon \in [\epsilon_0, \epsilon_0]$

$$\|y(\cdot, t) - \epsilon^2 \psi(\cdot, t)\|_s \leq C \epsilon^{7/2}$$

for  $t \in [\epsilon_0, T_0 \epsilon^{-1}]$   $C = C(\epsilon, \epsilon_0) \neq C = C(T_0)$   
 $\epsilon_0 = \epsilon_0(T_0)$  as well

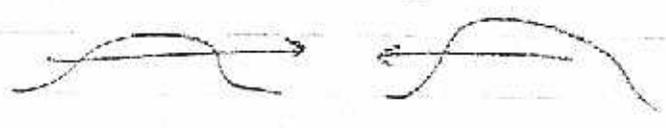
Sol what are the predictions?

- ① WW have a solitary wave  
( $A_L = \text{solitary wave}$ )  
 $A_R = 0$



- ② WW have a head on collision

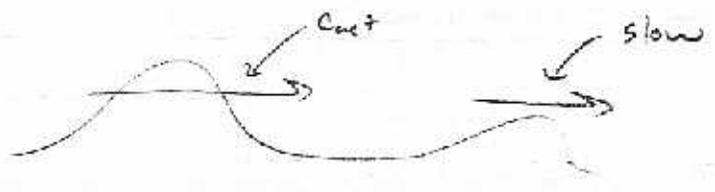
$A_L = \text{soliton}$  ,  $A_R = \text{soliton}$

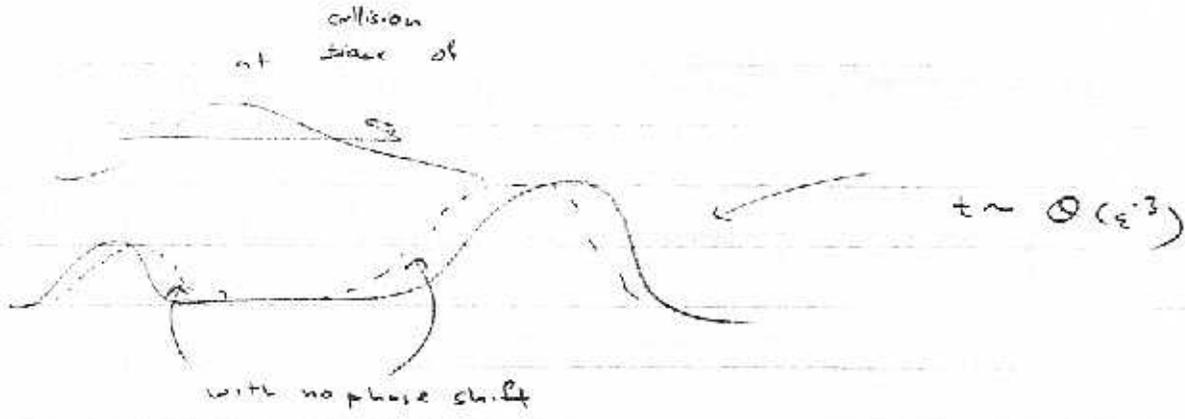


This is "fast" collision  
 $t \sim O(1)$

At collision, Waves add in Super position

- ③ Things akin to Multiple Solitons exist in water waves

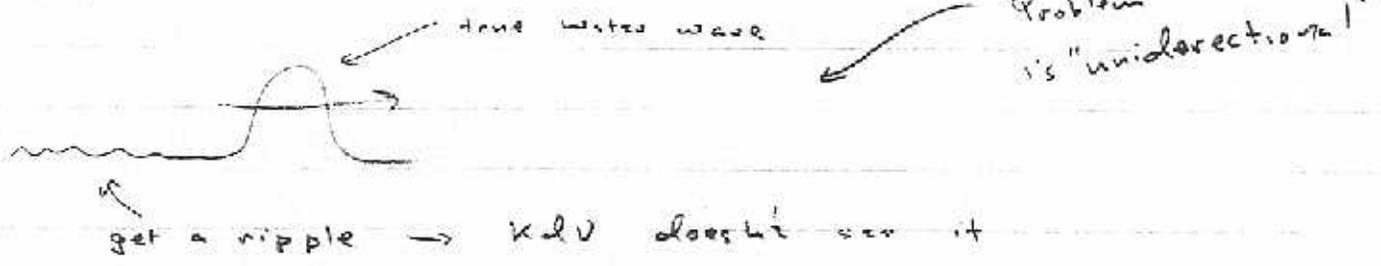




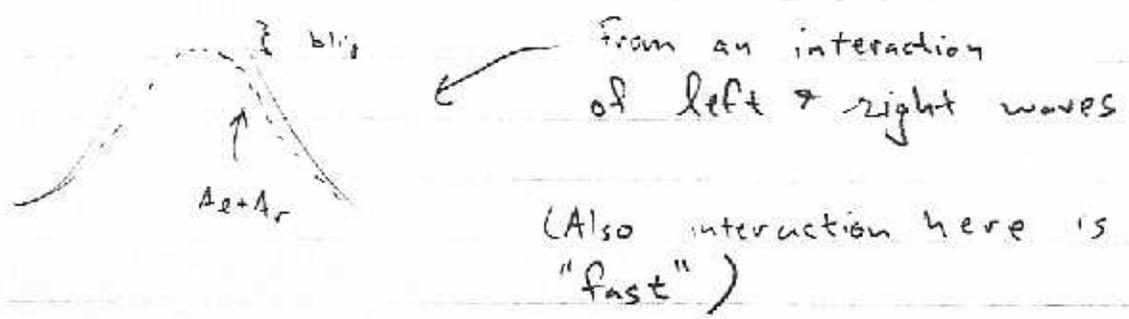
Remark: To see this sort of behavior  $T_0$  may need to be large  $T_0 \sim 10$ , which may make  $C$  large &  $\epsilon_0$  small

What is it so good?

① Dispersive Waves



② At time of collision in a head on collision height is slightly more than linear superposition



② Phase Shift isn't predicted very well

My Thesis  $\rightsquigarrow$  Fix it up!

How: keep KdV, but add stuff  $\rightarrow$  there are other ways!

Ansatz  $\rightsquigarrow$

same as before "index factor"

$$\psi(x,t) \approx \varepsilon^2 (A_L(\varepsilon(x+t), \varepsilon^3 t) + A_R(\varepsilon(x-t), \varepsilon^3 t)) + \varepsilon^4 F(A_L, A_R) + \varepsilon^4 (P(\varepsilon x, \varepsilon t) + B_L(\varepsilon(x+t), \varepsilon^3 t) + B_R(\varepsilon(x-t), \varepsilon^3 t))$$

Fix interaction problem

Fix "unidirectional" problems

Call this RHS  $\varepsilon^4 \frac{1}{2}$

Grinding through the formal computation, one has

$$i) \partial_x^2 P - \partial_\beta^2 P = 3 \partial_\beta^2 (A_L(\dots) A_R(\dots))$$

$$\begin{cases} -2\partial_t B_r = \frac{1}{3} \partial_\beta^3 B_r + \frac{3}{2} \partial_\beta (A_r B_r) + J_r \\ 2\partial_t B_l = \frac{1}{3} \partial_\beta^3 B_l + \frac{3}{2} \partial_\beta (A_l B_l) + J_l \end{cases}$$

we have  $\partial_\beta$  on  $A_r, P$  & their derivatives

(IW) is solvable easily & explicitly via (say) method of characteristics. If  $A_L, A_R$  are "spatially localized" (like a soliton), then  $\Psi$  is bounded over the long time scale

(Lxalv) are also explicitly Solvable (R. Sachs, Sotinger & Haragus-Courcelle) Though this is a night-marish business. They are easy to compute numerically.

① Problem  $\rightarrow$  Solus to these equation grow with time  $\rightarrow \|A_L(\cdot, T)\| \leq CT e^{CT}$

② But who cares,  $T = \epsilon^2 t$  is  $\mathcal{O}(1)$  on the long time scale

This shows  $\|y(\cdot, t) - \epsilon^2 z_2^y(\cdot, t)\|_S \leq C \epsilon^{1/2}$   
for  $t \in [0, T_0 \epsilon^3]$

PS! Same as before ...

Picture Time

- Further
- ① Improves accuracy, how to improve time scale?
  - ② Global in time solus like multi-solitons? (or really any "unidirectional" solus?)